

## **Timing, cause and consequences of mid-Holocene climate transition in the Arabian Sea**

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### **ABSTRACT**

We reconstruct centennial scale quantitative changes in surface seawater temperature (SST), evaporation-precipitation (from Mg/Ca and  $\delta^{18}\text{O}$  of surface dwelling planktic foraminifera), productivity (from relative abundance of *Globigerina bulloides*), carbon burial (from %CaCO<sub>3</sub> and organic carbon [%C<sub>org</sub>]) and dissolved oxygen at sediment-water interface, covering the entire Holocene, from a core collected from the eastern Arabian Sea. This defines the timing, consequences and possible causes of the mid-Holocene climate transition (MHCT). A distinct shift in evaporation-precipitation (E-P) is observed at 6.4 ka, accompanied by a net cooling of SST. The shift in SST and E-P is synchronous with a change in surface productivity. A concurrent decrease is also noted in both the planktic foraminiferal abundance and coarse sediment fraction. A shift in carbon burial, as inferred from both the %CaCO<sub>3</sub> and %C<sub>org</sub>, coincides with a change in surface productivity. A simultaneous decrease in dissolved oxygen at the sediment-water interface, suggests that changes affected both the surface and subsurface water. A similar concomitant change is also observed in other cores from the Arabian Sea as well as terrestrial records, suggesting a widespread regional MHCT. The MHCT coincides with decreasing low-latitude summer insolation, perturbations in total solar intensity and an increase in atmospheric CO<sub>2</sub>.

**Keywords:** Holocene, Climate, Transition, Temperature, Precipitation, Arabian

## **Introduction**

Early research suggested that the Holocene was an epoch of relative stable climate due to limited changes in climatic forcing, which apparently helped facilitated the rise of civilizations. Recent high-resolution centennial records, however, reveal distinct climatic events in the Indian subcontinent and adjoining seas during the Holocene (Staubwasser et al., 2003; Gupta et al., 2005; Dykoski et al., 2005; Yadav et al., 2011). Such climatic changes during times of limited natural boundary conditions can help in the understanding of Earth's response to increasing anthropogenic influences (Schewe and Levermann, 2012; Cook et al., 2015). The mid-Holocene climatic transition (MHCT) is a very prominent feature of the present interglacial (deMenocal et al., 2000; Weldeab et al., 2005; Roberts et al., 2011; Dixit et al., 2014; Vincenzo and Massimo, 2015). Defining the timing of the MHCT is important as it coincides with distinct changes in human civilizations and has been linked with anthropogenic activities (Ruddiman and Ellis, 2009; Kelly et al., 2013). Previous studies show that the MHCT in the Arabian Sea dates to between 7 to 5 ka (Sarkar et al., 2000; Govil and Naidu, 2010; Singh et al., 2011; Kessarkar et al., 2013; Naik et al., 2014). A centennial-scale multi-proxy record covering the entire Holocene can help better define the timing of the MHCT. A precisely dated MHCT can then be used to help understand the lead-lag relationship between different processes during major climate transitions in the eastern Arabian Sea. Such a record can further help to understand land-ocean interaction and to compare regional climatic events in the context of various global climatic changes. In this paper, we develop a record of mid-Holocene climate changes using a gravity sediment core (SK237 GC04, hereafter referred as Malabar core) from the continental slope of southeastern Arabian Sea (10°58.65' N, 74°59.96' E, water depth 1245 m). The core had centennial scale resolution and covered the last glacial interglacial transition (Fig. 1). The stable isotopic and trace element analysis of surface dwelling planktonic foraminifera from this core was used to define the timing of deglacial warming and precipitation-evaporation changes during the last glacial-interglacial transition (Saraswat et al., 2013). We use multi-proxy analysis on the Holocene section of this core to reconstruct both the surface and sediment-water interface conditions at centennial-scale resolution, throughout the Holocene, to define the timing, possible cause and consequences of the MHCT in the eastern Arabian Sea.

## **Oceanographic setting**

The eastern Arabian Sea is characterized by a seasonal reversal of surface currents in response to winds (Shankar *et al.*, 2002). The equatorward West India Coastal Current brings relatively more saline water to the eastern Arabian Sea prior to and during the southwest monsoon, whereas the

poleward winter monsoon current transports low salinity water from the western Bay of Bengal into the eastern Arabian Sea during boreal winter (Prasanna Kumar *et al.*, 2004). Increased productivity in this region is observed during the summer monsoon season (Levy *et al.* (2007; Fig. 1). The sea surface temperature (SST) in the eastern Arabian Sea increases during the pre-summer monsoon season, leading to the development of a small warm pool that peaks in April. The eastern Arabian Sea warm pool dissipates prior to the onset of summer monsoon precipitation as a result of upwelling (Shenoy *et al.*, 1999). The region is also marked by both seasonal shallow water hypoxia zone (Naqvi *et al.*, 2000) as well as a perennial intermediate water-depth oxygen minimum zone (OMZ) that extends from ~150 m to ~1200 m (Naqvi *et al.*, 1991; 2003).

## **Materials and methodology**

The Holocene section (top 100 cm, subsampled at 1 cm intervals) of the Malabar Core was used to reconstruct centennial-scale changes, which was dated using five Accelerator Mass Spectrometer  $^{14}\text{C}$  ages on mixed planktic foraminifera measured at the Center for Applied Isotope Studies, the University of Georgia, USA (Table 1; Saraswat *et al.*, 2013). Marine09 (Riener *et al.*, 2009) and Calib6.0 version (Stuiver and Riener, 1993) were used to calibrate the  $^{14}\text{C}$  ages. A reservoir correction ( $\Delta R$ ) of  $138 \pm 68$  yr for the eastern Arabian Sea was applied (Southon *et al.*, 2002). The age model utilizes the calibrated ages as tie-points and assumes linear sedimentation rate (Fig. 2). A minimum of 300 benthic foraminifera were picked and examined from each sample of the coarse sediment fraction ( $>63 \mu\text{m}$ ). The relative abundance of angular asymmetrical benthic foraminifera (AABF) was counted following the methods of Nigam *et al.* (1992; 2007). Planktic foraminifera were picked from the  $>125 \mu\text{m}$  sediment fraction. Changes in relative abundance of planktic foraminifera *Globigerina bulloides* (an indicator of high productivity), coarse sediment fraction (CF), organic carbon weight percentage ( $\%C_{\text{org}}$ ) and calcium carbonate weight percentage ( $\%CaCO_3$ ) were used to reconstruct past monsoon strength and associated changes (Fig. 3). The total  $\%C_{\text{org}}$  was measured with a CNS analyzer and inorganic carbon was measured with a coulometer. The precision of total carbon measurements was better than  $\pm 0.49\%$ . The precision of total inorganic carbon measurements, based on repeat analysis of laboratory standard after every ten samples, was better than  $\pm 0.22\%$ . The  $\%CaCO_3$  was calculated from the total inorganic carbon. Total  $\%C_{\text{org}}$  was calculated by subtracting total inorganic carbon from the total carbon. The cumulative error associated with  $\%C_{\text{org}}$  estimates was  $\pm 0.54\%$ . Seawater paleotemperature was reconstructed from changes in the Mg/Ca ratio of the white variety of planktic foraminifera *Globigerinoides ruber* (Saraswat *et al.*, 2013), which has proven a very robust method to quantify paleotemperatures (Lea *et*

al., 1999; Saraswat et al., 2005; Weldeab et al., 2005). The error in seawater temperature estimated from *G. ruber* Mg/Ca is  $\pm 1.3$  °C. The cumulative error was calculated from the standard deviation of replicate analyses of the samples, when available, combined with the standard deviation of consistency standards during the sample run, as well as the error associated with the calibration equation. The Malabar core data was compared with other records published from the Arabian Sea as well as Indian subcontinent (Fig. 4). Additionally, local evaporation-precipitation changes have been reconstructed by subtracting global ice volume contribution from temperature corrected stable oxygen isotopic ratio ( $\delta^{18}\text{O}_{\text{sw-ivc}}$ ) of *G. ruber* (Fig. 5). The error in  $\delta^{18}\text{O}_{\text{sw-ivc}}$  is assumed to be at least  $\pm 0.3\text{‰}$  because the error associated with  $\delta^{18}\text{O}_{\text{sw}}$  is  $\pm 0.3\text{‰}$  (Saraswat et al., 2012, 2013). The long-term orbital configuration driven low latitude insolation changes in mid-July at 10°N are from Lasker et al. (2004), while the total solar irradiance changes are from the Steinhilber et al. (2012). The Malabar core record was compared with the high-resolution lake sediment, speleothem and ice-core records, reconstructed from Asia, to test whether the mid-Holocene transition is confined only to the marine records or is also present in terrestrial Holocene records, (Figure 4) (Enzel et al., 1999; Dykoski et al., 2005; Sinha et al., 2006; Thompson et al., 2012).

## Results

Reconstructed Mg/Ca SST varies from 27.3 to 29.5°C from 2.9 to 11.0 ka. After an initial cooling in SST during the early Holocene, the SST was stable from 9.3 until 6.0 ka (Fig. 3). A net drop in SST is observed during the late Holocene, beginning at 6.0 ka. The most depleted  $\delta^{18}\text{O}_{\text{sw-ivc}}$  (0.14‰) is evident at 3.6 ka while it is the most enriched (1.19‰) during early Holocene (11.2 ka). A large drop in  $\delta^{18}\text{O}_{\text{sw-ivc}}$  from the most enriched value (1.19‰) at 11.2 ka to 0.29‰ at 10.6 ka, is evident during the early Holocene immediately after the MIS2/1 transition. The  $\delta^{18}\text{O}_{\text{sw-ivc}}$  is relatively stable between 9.9 and 6.8 ka. A net decrease with minor fluctuation in  $\delta^{18}\text{O}_{\text{sw-ivc}}$  is also evident between 6.8 and 6.4 ka. Subsequently a net increase in  $\delta^{18}\text{O}_{\text{sw-ivc}}$  is evident until 5.3 ka (Fig. 5). Another prominent decrease in  $\delta^{18}\text{O}_{\text{sw-ivc}}$  begins at 4.4 ka and continues until 3.6 ka. The %CaCO<sub>3</sub> increases throughout the early Holocene to 6.2 ka, with no further increase during the late Holocene. The lowest %CaCO<sub>3</sub> (8.8%) is reported at 11.3 ka whereas the highest (25.8%) is at 3.2 ka. A clear increase in relative abundance of *G. bulloides* is evident during this same interval. The minimum *G. bulloides* relative abundance (3.3%) occurred during the early Holocene (10.0 ka), while the maximum abundance (22.0%) occurred during the late Holocene (0.6 ka). A net increase in %C<sub>org</sub> is evident during the early Holocene, followed with nearly uniform values between 6.1 and 4.1 ka. The %C<sub>org</sub> varied from 1.5 to 5.3% at 11.2 and 0.29 ka, respectively. The coarse fraction varied from

1.36% to 4.98% at 8.6 and 4.5 ka, respectively. The continuous gradual increase in coarse fraction during the early Holocene stops at 6.5 ka. The relative abundance of AABF gradually decreased during the early Holocene until 6.3 ka. A gradual increase in AABF is evident during the late Holocene (Fig. 3).

## Discussion

The residual  $\delta^{18}\text{O}_{\text{sw-ivc}}$ , obtained by subtracting global ice volume contribution from temperature corrected  $\delta^{18}\text{O}$ , is a proxy for seawater salinity (Weldeab et al., 2005; Govil and Naidu, 2010; Saraswat et al., 2012; Kessarkar et al., 2013). The salinity at the core site is a function of evaporation-precipitation budget. Additionally, the transport of water from the Bay of Bengal to the southeastern Arabian Sea also affects the salinity in this region (Prasanna Kumar et al., 2004). Therefore, the most enriched  $\delta^{18}\text{O}_{\text{sw-ivc}}$  during the Marine Isotopic Stage 2/1 transition, suggests an extremely weak monsoon at this time. The reduced transport of low salinity water from the Bay of Bengal during early Holocene, as inferred from relatively low Ba/Ca ratio in surface dwelling planktic foraminifera (Saraswat et al., 2013), further enriched  $\delta^{18}\text{O}_{\text{sw-ivc}}$  in the southeastern Arabian Sea (Fig. 5). The depleting  $\delta^{18}\text{O}_{\text{sw-ivc}}$  beginning at 11.2 ka and continuing until 6.4 ka suggests a progressive strengthening of the monsoon. The short initial abrupt phase of this long gradually strengthening early Holocene monsoon change was driven by vigorous transport of low salinity water from the Bay of Bengal, as is evident from a marked increase in Ba/Ca ratio (Saraswat et al., 2013). The temperature during the short abrupt phase of early Holocene monsoon strengthening was relatively warm and stable. A prominent strengthening of monsoon is also evident between 6.8 and 6.4 ka, when the  $\delta^{18}\text{O}_{\text{sw-ivc}}$  decreased by  $\sim 0.3\text{‰}$ , followed by an equally weakened monsoon until 5.3 ka (Fig. 5). The SST was relatively stable during the larger part (between 9.3 and 6.0 ka) of the early Holocene phase of monsoon strengthening. An early Holocene weak monsoon phase has also been reported from Rajasthan in western India (Singh et al., 1974). The subsequent strengthening of southwest monsoon at 9.6 and 8.6 ka is attributed to changing solar radiation (Sirocko et al., 1993). In archaeological records this time coincides with the emergence of Neolithic tools during the early Holocene. The Neolithic phase is characterized by planned habitation termed early village settlement (Bellwood, 1996). Another important aspect of this phase of human development was planned food production, domestication of animals and, at later stages, introduction of pottery, which was probably supported by the onset of a stable monsoon phase (Bellwood, 1996).

A distinct change in both the geochemical and faunal proxies is observed during the mid-Holocene in the Malabar core, suggesting a climatic event that affected the southeastern Arabian Sea. The

progressive strengthening of monsoon since the early Holocene, as inferred from depleting  $\delta^{18}\text{O}_{\text{sw-ivc}}$ , continued until 6.4 ka, followed by an increase in  $\delta^{18}\text{O}_{\text{sw-ivc}}$  until 5.3 ka. The enriched  $\delta^{18}\text{O}_{\text{sw-ivc}}$ , post mid-Holocene transition, suggests either reduced monsoon influence and/or reduced input of low salinity water from the Bay of Bengal. The transition in  $\delta^{18}\text{O}_{\text{sw-ivc}}$  is synchronous with a shift in surface productivity, decrease in both the planktic foraminiferal abundance as well as coarse-fraction abundance, suggesting a decrease in planktic foraminiferal population. The decreased surface productivity combined with the decrease in planktic foraminiferal population supports reduced monsoon influence during the post mid-Holocene transition. The changes in  $\% \text{CaCO}_3$  and  $\% \text{C}_{\text{org}}$  are also synchronous suggesting a concomitant response of carbon burial to changes in evaporation-precipitation and productivity. The gradual increase in dissolved oxygen concentration at the sediment-water interface, until 6.3 ka, inferred from the decrease in relative abundance of angular asymmetrical benthic foraminifera, also stopped and it began to decrease after the mid-Holocene (Fig. 3). The increase in relative abundance of AABF, post mid-Holocene, suggests decrease in dissolved oxygen at the sediment-water interface.

The mid-Holocene changes in the Malabar core proxy data are also evident in stable oxygen isotopic record in other cores published from the Arabian Sea (Staubwasser et al., 2003; Anand et al., 2008). A synchronous change in  $\delta^{18}\text{O}_{\text{ruber}}$  is observed in core SK17 collected from the central-eastern Arabian Sea, from a comparable water depth (Anand et al., 2008). Similarly, a distinct shift towards depleted  $\delta^{18}\text{O}_{\text{ruber}}$  is also evident in core 63KA-41KL (Staubwasser et al., 2003). Interestingly, a distinct shift in denitrification, as inferred from changes in  $\delta^{15}\text{N}$ , is also evident at the same time in core RC27-23 collected from the western Arabian Sea and is synchronous with the changes in relative abundance of AABF in the Malabar core (Altabet et al., 2002; Fig. 4). A distinct change in productivity at the same time was also inferred from Ocean Drilling Program (ODP) Hole 723A, Leg 117, off Oman, northwestern Arabian Sea (Gupta et al., 2005). Several episodes of change in productivity during the Holocene were inferred at this site, as it lies at the core of summer monsoon induced upwelling. A comparison of the Malabar core data with other cores from the southeastern Arabian Sea, as well as the Bay of Bengal, further suggests a similar change in both the  $\% \text{CaCO}_3$ , and  $\% \text{C}_{\text{org}}$  in many of these cores, even though the magnitude and timing is different (Naidu and Malmgren, 1995; Agnihotri et al., 2003; Chauhan, 2003; Anil Kumar et al., 2005; Narayana et al., 2009; Singh et al., 2011; Ahmad et al., 2012; Naik et al., 2014). The difference in magnitude and changes in the proxies, in various marine cores, is attributed to the difference in regional oceanographic features, mainly productivity, surface current direction, terrigenous input and dissolved oxygen, which affect both the flux of both the autochthonous and allochthonous material,

and its deposition and diagenetic alteration. A few of the cores were collected from the oxygen-minimum zone, which affects the carbonate preservation, while the others are from the regions outside the oxygen-minimum zone (Canfield and Raiswell, 1991; Levin, 2003). A part of the difference in timing of transition in the cores from different parts of the Arabian Sea, is also attributed to the coarse sample resolution (usually multi-centennial to a few ka) in most of the cores (Naidu and Malmgren, 1995; Agnihotri et al., 2003; Chauhan, 2003; Anil Kumar et al., 2005; Narayana et al., 2009; Singh et al., 2011; Ahmad et al., 2012; Naik et al., 2014), which obscures a clear comparison. In spite of the differences in physico-chemical environment prevailing at the coring sites, the commonality of the mid Holocene transition in cores collected from the southeastern Arabian Sea, suggests that it is a regional feature.

A major shift towards highly depleted  $\delta^{18}\text{O}$  is also evident during the mid-Holocene, in Guliya ice-core record (Thompson et al., 2012). A similar abrupt change to relatively drier condition, beginning at  $\sim 6.1$ - $5.4$  ka, thus coinciding exactly with the mid-Holocene transition in the Malabar core, has been inferred from C/N and  $\delta^{13}\text{C}$  analysis of a core collected from Nal Sarovar, a lake in the western India (Prasad et al., 1997). Similarly, a substantial decrease in summer rainfall, between 6.5 and 4.5 ka, is also evident from the pollen and grain-size analysis of a sediment core collected from the Lunkaransar saltwater lake (Swain et al., 1983; Enzel et al., 1999). A much larger change in summer monsoon rainfall in this core was, however, evident at a later age. The mid Holocene transition also coincides with an abrupt increase in number of artifacts recovered from the Indus valley region as well as the simulated population size (Lemman and Khan, 2012). Therefore, a distinct change during the mid-Holocene is apparent in both the marine as well as terrestrial records, suggesting a widespread climate transition during the mid-Holocene in the Indian subcontinent and adjoining seas.

A prominent climatic transition at 4.2 ka has earlier been reported in a core collected from the shallow marine region off Pakistan, which is affected by Indus influx (Staubwasser et al., 2003). This 4.2 ka event was suggested to be responsible for the major shift in the Harappan settlements in western India. In the Malabar core, a substantial cooling from  $28.8$  °C to  $27.7$  °C from 4.8 to 4.2 ka as well as depleted  $\delta^{18}\text{O}_{\text{sw-ivc}}$  from 0.46‰ to 0.14‰ at 4.4 ka to 3.6 ka, is evident, suggesting a widespread transition. The depleted  $\delta^{18}\text{O}_{\text{sw-ivc}}$  across the 4.2 ka event, suggests intensification of monsoon. However, the previous studies suggest drier conditions after 4.2 ka (Staubwasser et al., 2003). Similar dry conditions after 4 ka, following a phase of intense monsoon between 10 ka and 4 ka, are apparent in the Badain Jaran Desert in western Inner Mongolia region of northern China (Yang et al., 2011, 2013). The increased transport of low salinity water from the Bay of Bengal is an

alternate source for the depleted  $\delta^{18}\text{O}_{\text{sw-ivc}}$  after 4.2 ka. As cooling accompanied the 4.2 ka event, it is highly likely that the winter monsoon current was stronger, increasing the transport of low salinity water into the southeastern Arabian Sea. The lack of Ba/Ca data in the upper part of the core does not allow to examine the relative contribution of winter monsoon current. Mixed signatures of this 4.2 ka event have also been reported from other terrestrial and marine records (MacDonald, 2011). A short-lived inundation was reported at 4.2 ka in the Lunransar Lake (Enzel et al., 1999). The Nal Sarovar lake level dropped at 4.2 ka, as inferred from a brief increase in C/N ratio to values characteristics of terrestrial vegetation (Prasad and Enzel, 2006). One, out of the several episodes of repeated decrease in relative abundance of *Globigerina bulloides* throughout the Holocene, in a core collected off Oman margin, that were inferred as an evidence for reduced upwelling and thus weaker monsoon winds, occurred at 4.2 ka (Gupta et al., 2003). A relook at the chronology of various phases of Harappan civilization suggests, however, that the 4.2 ka event does not lead to any immediate major change. The mature phase of the Indus Civilization began at 4.5 ka and continued till 3.9 ka (2500-1900 BC; Possehl, 1999).

#### *Comparison with global records*

A synchronous MHCT, is also noted in the African climate, as inferred from the multi-proxy analysis of Hole 658C drilled off Cap Blanc, Mauritania (deMenocal et al., 2000). A prominent shift from cooler to warmer temperature is observed between  $\sim 6$  ka and  $\sim 5$  ka, thus matching with the mid-Holocene transition noted in the Malabar core. A distinct change in precipitation is evident in the Mediterranean region beginning at  $\sim 6$  ka (Roberts et al., 2011). Widening of tropical dry Saharan belt that abruptly terminated the Mediterranean cyclogenesis and African Humid period was also dated between 5.5 and 4.5 ka (Vincenzo and Massimo, 2015). An abrupt temperature fluctuation accompanied by increase in local salinity between 6.5 and 5.7 ka, suggesting a weaker monsoon, has also been reported from the eastern Atlantic (Weldeab et al., 2005). An abrupt change in *G. bulloides* relative abundance, which matches very well with the mid-Holocene transition in Malabar core, is inferred as a shift in winds associated with summer monsoon (Gupta et al., 2005). A gradual decline in sedimentary  $\delta^{15}\text{N}$  was reported between 7 and 3 ka from the oligotrophic equatorial Pacific (Kienast et al., 2008). The MHCT reported in the Arabian Sea is later than the global early to mid-Holocene climatic transition (dated between 8.2 and 7.8 ka) reported from both the equator and poles (Stager and Mayewski, 1997). A concomitant change in both the eastern and western Arabian Sea as well as the eastern Atlantic, suggests that a common mechanism controlled climatic conditions in the tropical regions. We suggest that this mid-Holocene climate shift primed the conditions for the rise and subsequent proliferation of several advance civilizations throughout the world during the late



Holocene, including the Harappan Civilization (Madella and Fuller, 2006). Therefore, the mid-Holocene climate shift was an important event, which may have had a profound influence on the development of human settlements during the Holocene.

#### *Causes and consequences*

The MHCT, as evident in the eastern Arabian Sea, can be brought by a change in the tropical processes, including the monsoon. Monsoon is basically driven by latitudinal position of inter-tropical-convergence-zone (ITCZ) combined with the land-ocean pressure gradient at present as well as in the past (Haug et al., 2001; Gadgil, 2003; Fleitmann et al., 2007). Both the latitudinal position and seasonal shifts in ITCZ, as well as the land ocean pressure gradient, primarily depend on temperature. As the insolation influences global temperature, the Malabar core data was compared with long-term trends in the low latitude July insolation (Laskar et al., 2004). Although the mid-Holocene transition evident in marine and terrestrial records from eastern Arabian Sea and Indian subcontinent, coincides with the decreasing phase of low-latitude insolation, a decrease in July insolation began much earlier (8 ka) than the transition evident in the Malabar core (Fig. 5). In addition to the long-term orbital configuration induced insolation changes, solar activity causes short-term frequent variation in the total solar irradiance (TSI). Changes in TSI ( $\Delta$ TSI) have widespread impact on Earth's climate, including the Asian monsoon (Gupta et al., 2005; Nigam et al., 1995; Wang et al., 2005). A comparison of the Malabar core record with the Holocene TSI records (Steinhilber et al., 2009) reveals that the mid-Holocene climatic transition coincides with large variation in  $\Delta$ TSI (Steinhilber et al., 2009; Fig. 5). However, several large-scale fluctuations are also noted in  $\Delta$ TSI, prior to and after the mid-Holocene transition. Atmospheric CO<sub>2</sub> concentration is also tightly coupled with global climate, especially temperature (Marcott et al., 2014). A prominent shift in atmospheric CO<sub>2</sub> concentration towards a continuously increasing trend is evident during the mid-Holocene. Therefore, we suggest that the combined effect of continuously decreasing summer insolation and other associated processes in combination with perturbations in TSI drove the MHCT.

As the Malabar core is located in the slope region, temporal changes in the oceanographic processes near the core site should also be seen in the context of long-term regional sea-level changes. The continental shelf gets exposed during the glacial drop in sea level and the subsequent interglacial rise in sea level, inundates the shelf. The exposure and inundation of the continental shelf during glacial-interglacial transitions, affects the surface circulation as well as the material flux pathways and its spatial extent. At present, a majority of the terrestrial runoff is restricted to the inner shelf (Chauhan et al., 2011; 2012) and during the times of low sea-stand, the same material will reach deeper

regions, thus affecting the flux and subsequent diagenetic processes. The entire continental shelf was exposed during the last glacial maximum, as a result of lowering of sea-level by ~120 m (Siddall et al., 2003; Grant et al., 2012). The subsequent inundation of the shelf, affected the spatial extent of terrigenous input as well as the coastal current intensity and direction in the southeastern Arabian Sea (Ramaswamy and Nair, 1989; Chauhan et al., 2000; 2011). A comparison of the core data with the regional as well as global sea-level curve, reveals a relative stabilization of the sea level only at ~6.4 ka (Hashimi et al., 1995; Siddall et al., 2003; Grant et al., 2012), suggesting that the sea level changes also contributed towards the mid-Holocene transition noted in the Malabar and other cores (Staubwasser et al., 2003; Anand et al., 2008; Naik et al., 2014) collected from the southeastern Arabian Sea.

The mid-Holocene change in local evaporation-precipitation is synchronous with the surface productivity, carbonate burial and bottom water dissolved oxygen, within age uncertainties. As perturbations in tropical rainfall are reflected in global methane records, the Malabar core data was compared with ice-core methane records for the Holocene. The signatures of this monsoon driven mid-Holocene transition are, however, not reflected in the ice-core methane records until ~5 ka (Loulerge et al., 2008; Fig. 5). Therefore, the mid-Holocene transition as recorded in the Malabar and other cores collected from the southeastern Arabian Sea, as well as in the terrestrial regions, was not strong enough to tweak the global methane concentration until ~5 ka. A synchronous change is, however, seen in atmospheric carbon-dioxide concentration as it continuously increases after ~6.5 ka (Monnin et al., 2001). The widespread nature of MHCT in Indian as well as other global records is intriguing. The possible reason for the widespread nature of MHCT is that at present as well as in the past, the monsoon in this region is driven by latitudinal position of inter-tropical-convergence-zone (ITCZ) (Haug et al., 2001; Gadgil, 2003; Fleitmann et al., 2007). We argue that ITCZ is the key driver responsible for the synchronous transition in not only African and Indian climate but also other global climate systems.

We suggest that this mid-Holocene climate shift primed the conditions for the rise and subsequent proliferation of several advanced civilizations of the world. The broad time bracket for the mature phase of the Indus Civilization is regarded to be 4.5 to 3.9 ka (2500-1900 BC; Possehl, 1999). Agrawal and Kusumgar (1974) proposed a time bracket for mature Harappan phase at 4.3 to 4.0 ka (2300-2000 BC) for nuclear region and 4.2 to 3.7 ka (2200-1700 BC) for peripheral zone of Harappan Civilization. Later, Allchin and Allchin (1982) also proposed the same chronology for the mature phase of the Indus civilization. Subsequently, many chalcolithic cultures flourished in several parts of south east Rajasthan, western MP, and Maharashtra during 4.0 to 3.0 ka (2000-1000 BC)

(Sankalia, 1968). The trade connection between Indus and Mesopotamian civilization also flourished during 2300 to 1900 BC (Gadd and Smith, 1924).

## **Conclusions**

Centennial scale change in seawater temperature, evaporation-precipitation, carbonate burial and upwelling induced productivity were reconstructed from the southeastern Arabian Sea using faunal and geochemical proxies. We report a major shift in proxies during the mid-Holocene (6.8 to 6.2 ka). A comparison of the Malabar core data with previously published records shows the regional nature of the MHCT. The MHCT is also evident in the terrestrial records. The widespread nature of the MHCT event in both the marine and terrestrial records suggests that the mid-Holocene transition in the Malabar core, affected the entire Asia and adjoining seas. A similar shift in the African climate, during the mid-Holocene, suggests that a common mechanism was responsible for the mid-Holocene transition in the tropical regions. We suggest that this MHCT primed the conditions for the rise and subsequent proliferation of several advanced civilizations of the world.

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## Figures

Figure 1 The core location in the southeastern Arabian Sea is marked by an open black rectangle. Other cores discussed in the text are marked by filled black star (1. SK17, Anand et al., 2008; 2. 63KA-41KL, Staubwasser et al., 2003; 3. RC27-23, Altabet et al., 2002). The template is surface primary productivity (chlorophyll-a, mg/m<sup>3</sup>) in the Arabian Sea during the southwest monsoon season (June-July-August-September) (Acker and Leptoukh, 2007).

Figure 2 The age-depth relationship used to establish the chronology of the Holocene section of the core SK237 GC04. The AMS ages were used as tie-points and ages for the intervals between two tie-points, were obtained by linear interpolation. Further details of the age model are discussed in Saraswat et al. (2013).

Figure 3 Changes in faunal and geochemical parameters in the Malabar core. A. Mg/Ca ratio in *G. ruber*, a proxy for seawater temperature; B.  $\delta^{18}\text{O}_{\text{ruber}}$  which has been used to determine  $\delta^{18}\text{O}_{\text{sw-ivc}}$ , a proxy for local precipitation-evaporation changes at the core site; C. CaCO<sub>3</sub> weight percentage; D. Organic carbon weight percentage (C<sub>org</sub>); E. Relative abundance of planktic foraminifer *G. bulloides*, a proxy for primary productivity; F. Percentage of coarse fraction (CF) (>63  $\mu\text{m}$ ); G. Abundance of all planktic foraminifera per gram dry sediment; and H. Relative abundance of angular asymmetrical benthic foraminifera (AABF), an indicator of bottom water oxygenation. The grey bar marks the mid-Holocene transition in various proxy indicators.

Figure 4 A comparison of Malabar core data (A) with representative marine and terrestrial records from the Arabian Sea and Asian subcontinent, namely SK17  $\delta^{18}\text{O}_{\text{ruber}}$  (Anand et al., 2008) (B); 63KA-41KL  $\delta^{18}\text{O}_{\text{ruber}}$  (Staubwasser et al., 2003) (C); ; Dongge Cave speleothem  $\delta^{18}\text{O}$  (Dykoski et al., 2005) (D); Guliya ice core  $\delta^{18}\text{O}$  (Thompson et al., 2012) (E) and RC27-23  $\delta^{15}\text{N}$  (Altabet et al., 2002) (F). The vertical grey bar marks the mid-Holocene transition in various proxy indicators in both the Malabar core and terrestrial records.

Figure 5 A comparison of Malabar core Mg/Ca SST (the black line is 3 point running average) (A), and  $\delta^{18}\text{O}_{\text{sw-ivc}}$  (B) with relative sea level (RSL) (1-ka moving Gaussian filter) (Grant et al., 2012) (C), atmospheric CO<sub>2</sub> (Monnin et al., 2001) (D), CH<sub>4</sub> from European Project for Ice Coring in Antarctica (EPICA) (Loulerge et al., 2008) (E), change in total solar irradiance ( $\Delta\text{TSI}$ ) (Steinhilber et al., 2009) (F) and low-latitude insolation (Laskar et al., 2004) (G) during the Holocene. The vertical grey bar marks the mid-Holocene transition in various proxy indicators in both the Malabar core and other records.

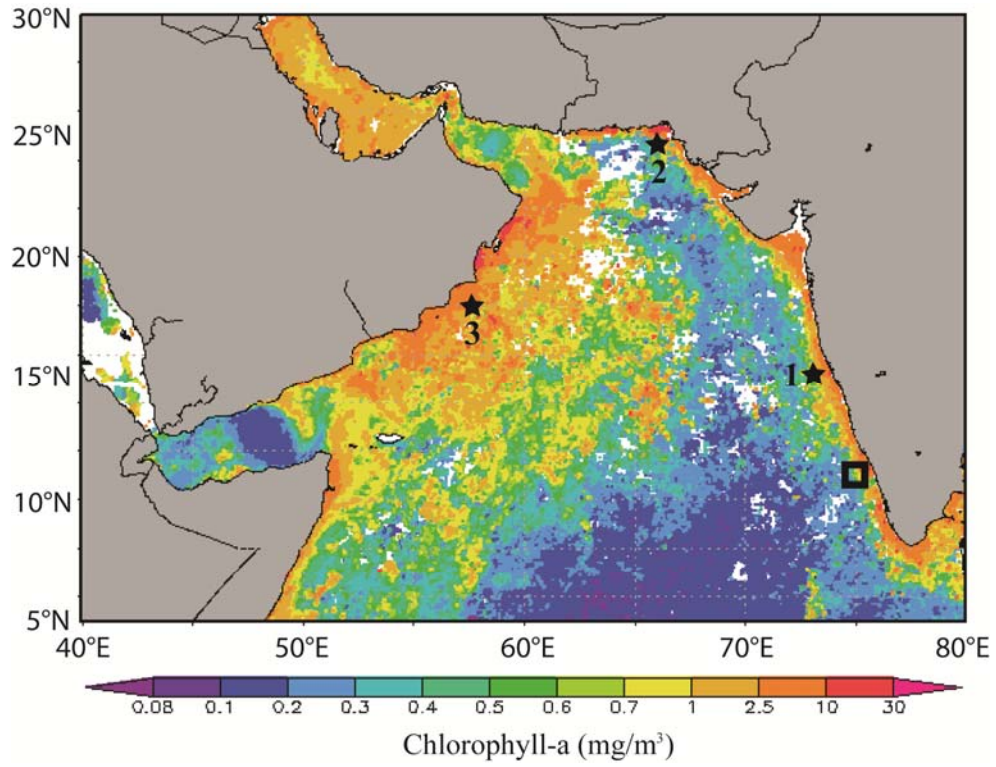


Figure 1

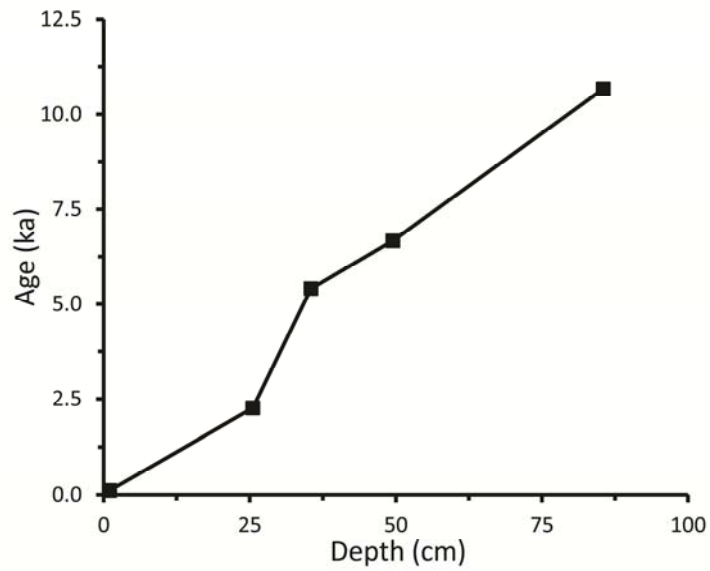


Figure 2

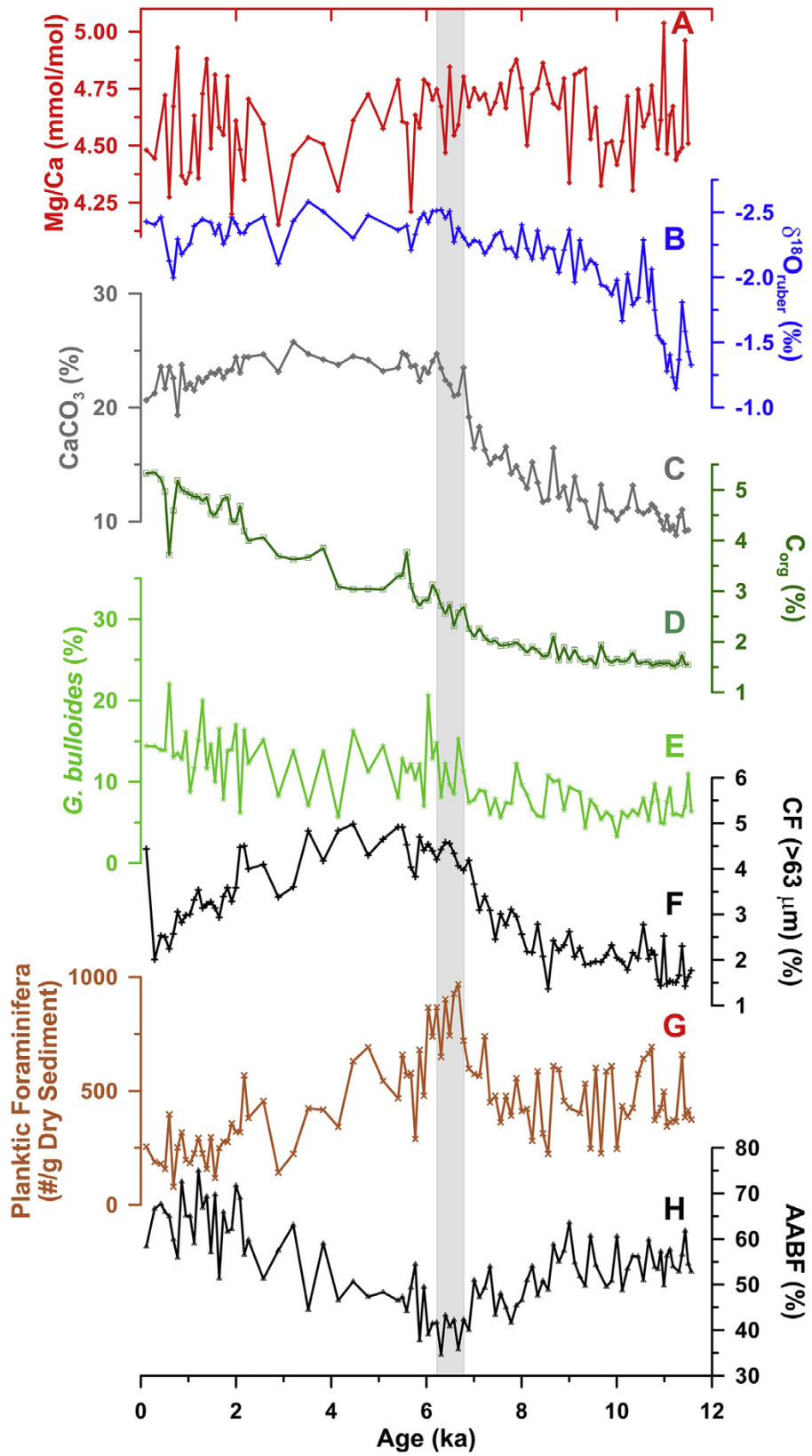


Figure 3

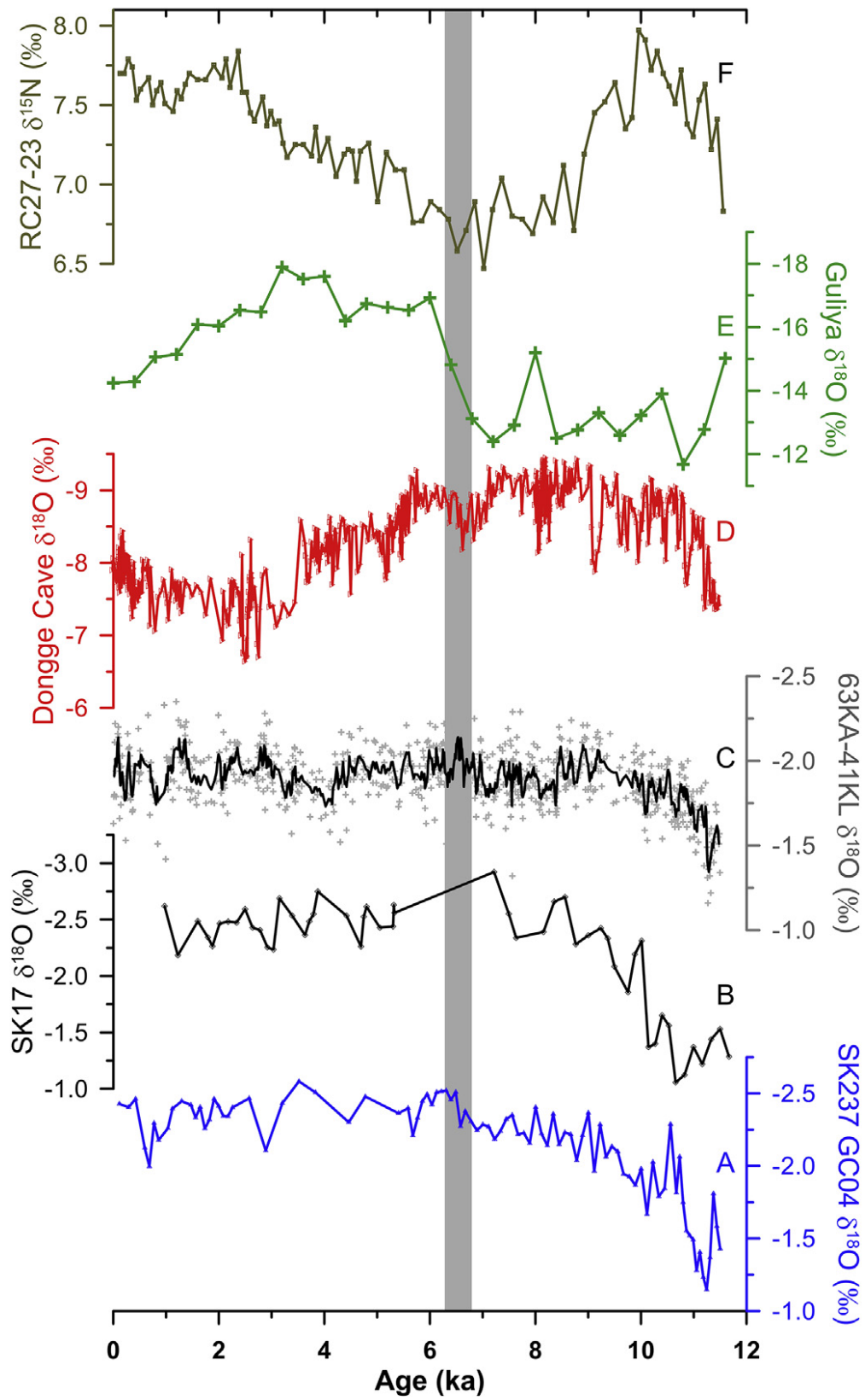


Figure 4

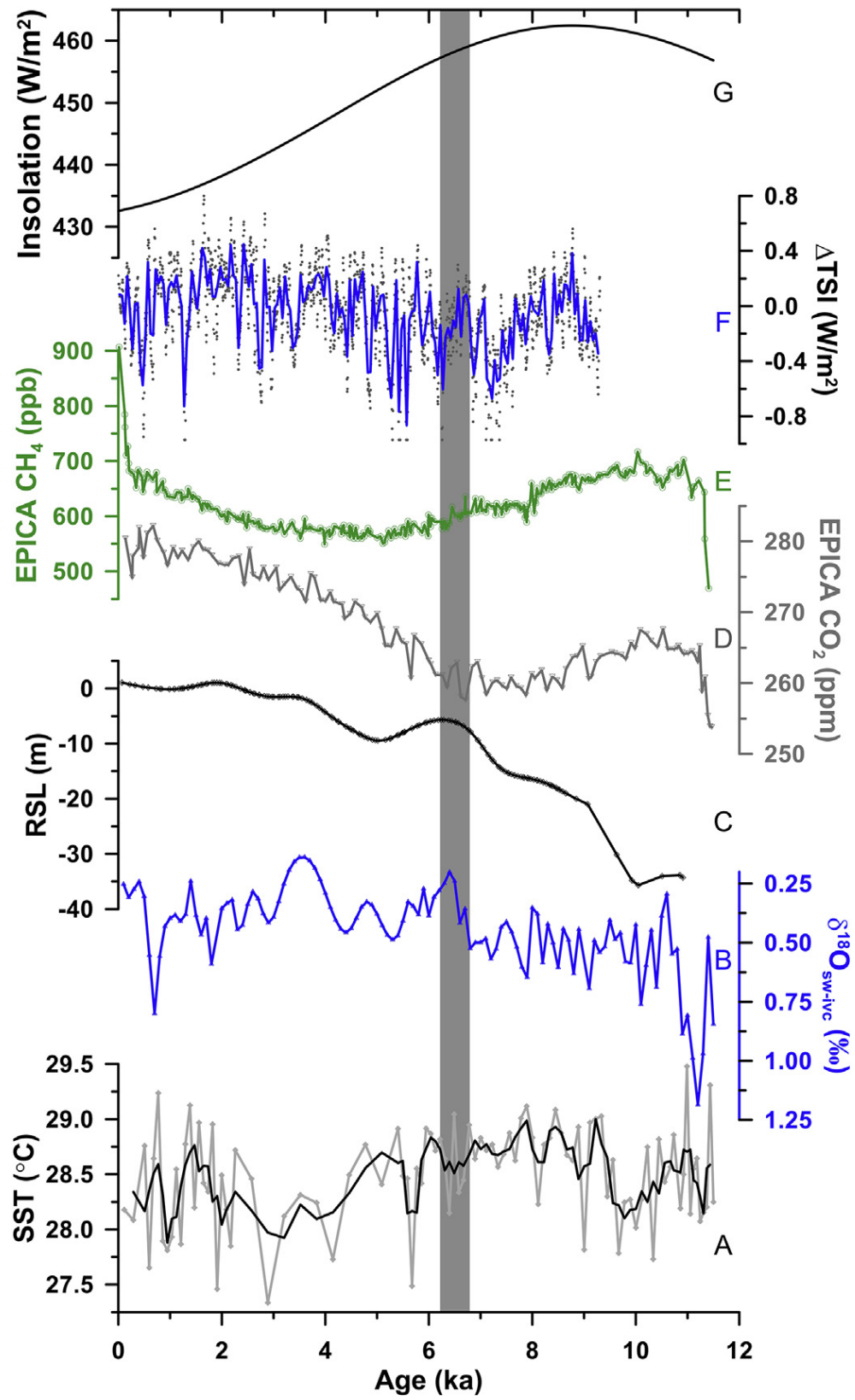


Figure 5

Table 1 Details of five accelerator mass spectrometer radiocarbon ages obtained on mixed planktic foraminifera measured at the Center for Applied Isotope Studies, the University of Georgia, USA.

<b>Lab #</b>	<b>Sample Depth (cm)</b>	<b><sup>14</sup>C Age (yr BP)</b>	<b><sup>14</sup>C Age Error (±)</b>	<b>Calib. Age- Range (1σ) (yr, BP)</b>	<b>Calib. Age- Range (2σ) (yr, BP)</b>	<b>Calib. Age (Median Probability) (yr, BP)</b>
UGAMS5378	0.5	620	25	0-146	0 - 246	115
UGAMS5739	25.5	2730	30	2163-2336	2075 - 2460	2260
UGAMS 23868	35.5	5190	30	5306 - 5477	5263 - 5574	5406
UGAMS5740	49.5	6360	30	6580-6760	6484 - 6846	6670
UGAMS5935	85.5	9930	30	10556-10749	10506 - 10940	10670